



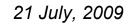
#### Mini-SAR

#### Initial Results from Calibration, Mapping, and Analysis



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NASA Lunar Science Forum







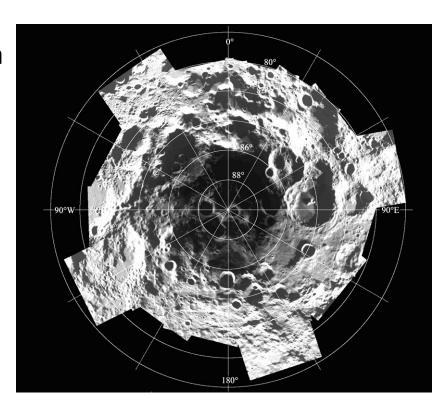
### **Experiment Objectives**

Map the deposits of both poles of the Moon (> 80° lat.) at optimum viewing angles (~40°) to characterize permanently dark areas and definitively determine their RF backscatter properties using both SAR and scatterometry

Complete the global map of the Moon by mapping dark regions in lunar polar areas

Characterize the physical nature of the polar regolith and surface

SAR mapping of other targets of opportunity as possible



Moon South Pole Clementine 750 nm base map





## **Circular Polarization Ratio (CPR)**



Ratio of received power in both right and left senses

Normal rocky planet surfaces = polarization inversion (receive opposite sense from that transmitted)

"Same sense" received indicates something unusual:

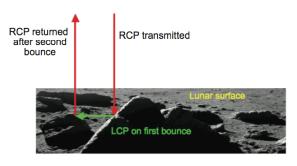
double- or even-multiplebounce reflections

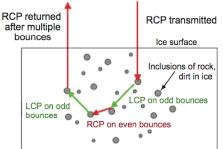
Volume scattering from RFtransparent material

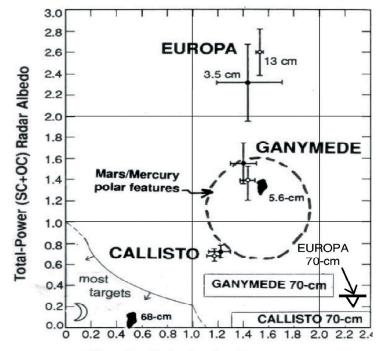
High CPR (enhanced "same sense" reception) is common for fresh, rough (at wavelength scale) targets and water ice

High CPR caused by surface roughness/scattering

High CPR caused by ice/volume scattering







Circular Polarization Ratio,  $\mu_c = SC/OC$ 







#### **Does Ice Exist On The Moon?**

#### Clementine

On one orbit (234) found evidence for high same-sense backscatter (high CPR; CBOE?) over dark areas near poles

Other orbits (e.g., 235; over sunlit areas) show no enhancement

#### **Lunar Prospector**

Neutron spectrometer detects "excess" hydrogen (~2-3x global average), but not phase state

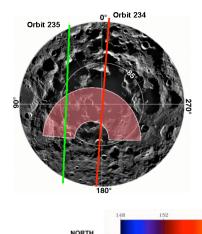
Enhanced H<sub>2</sub> over poles; consistent with ~1-2 % ice or excess retained solar wind

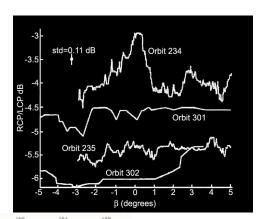
No H<sub>2</sub> signal in fast neutrons; upper ~10 cm of surface desiccated

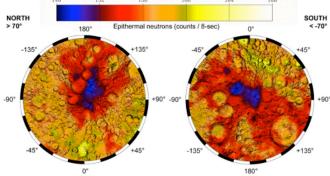
#### Earth-based radar

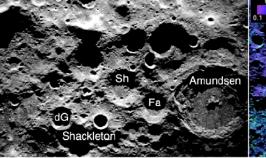
Patchy, high CPR found in both sunlight and polar darkness

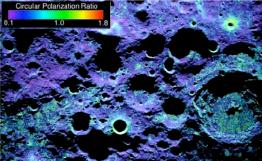
All surface roughness or two different scattering mechanisms?











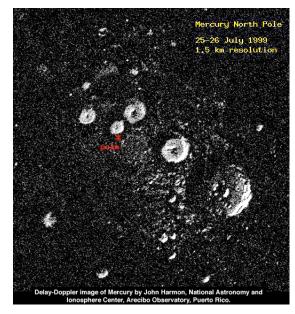


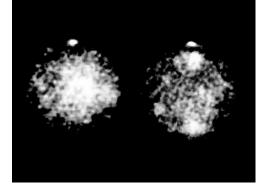


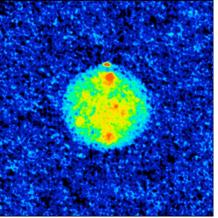


### What does high CPR mean?

- Is high CPR diagnostic of surface roughness or composition?
- Mercury shows high CPR within polar crater floors (Harmon et al., 2000)
- Mercury also shows high CPR areas at mid-latitudes (Muhleman, 1991)
- On Mercury, it is accepted that high CPR may have different causes, depending on local geological circumstances
- High CPR on the Moon may likewise have multiple causes











#### Mini-SAR



#### Imaging Radar on the Chandrayaan-1

Mini-SAR is an S-band (13 cm) imaging radar with hybrid polarity architecture

Map both polar regions at 75 m/pixel Transmit LCP, receive **H** and **V** linear, coherently

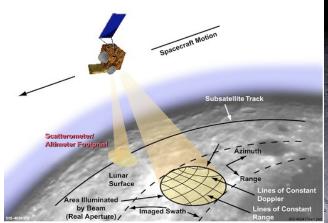
Use Stokes parameters and derived "daughter" products to describe backscattered field

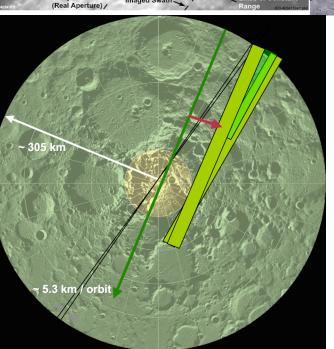
Map locations and extent of anomalous radar reflectivity

See polar dark areas (not visible from Earth)

Cross-correlate with other data sets (topography, thermal, neutron)

LRO version (Mini-RF) has two bands  $(\lambda=13 \text{ and } 5 \text{ cm})$ , high-resolution zoom mode (15 m/pixel)











#### **Mini-SAR Calibration Activities**

Pre-flight laboratory measurements; antenna pattern mapped

Translunar calibration (cancelled)

Arecibo radiotelescope calibration (31 Jan. 2009)

Ground transmit pure RCP signal to C-1

Map Rx channel gain/separation

Greenbank radiotelescope cal (27 Feb. 2009)

C-1 Tx to Earth

Measure Tx signal

Articulate spacecraft to map antenna pattern

Sample data and calibration data takes

SAR take 17 Nov. 2008 - both poles (first data)

SAR take 7 Jan. 2009 - both poles, including steep incidence pass

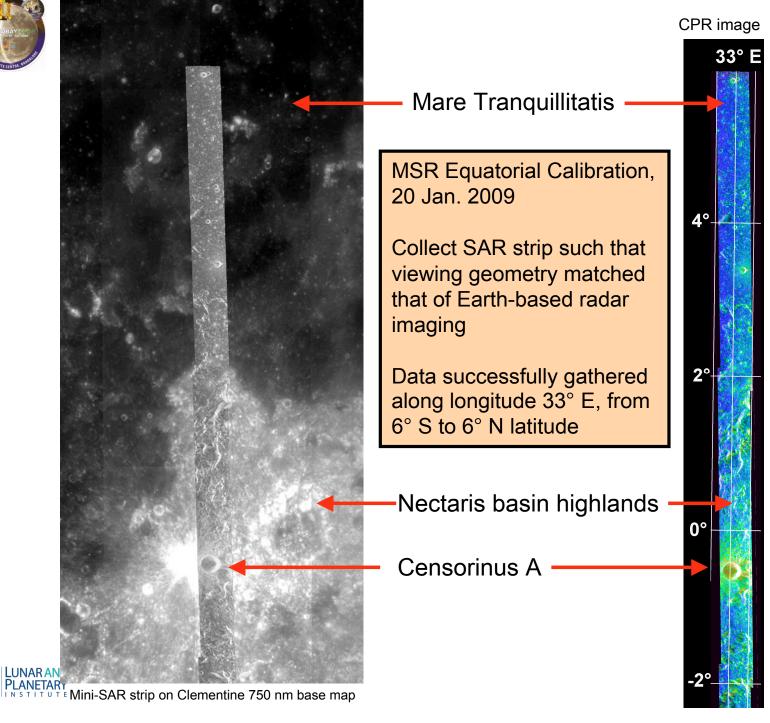
Scatterometry take 18 Jan. 2009 - 11° to 18° N along 54° E

SAR take 20 Jan. 2009 -- Equatorial (6°S to 6°N) along 32° E

Work to fully calibrate the Mini-SAR instrument is ongoing





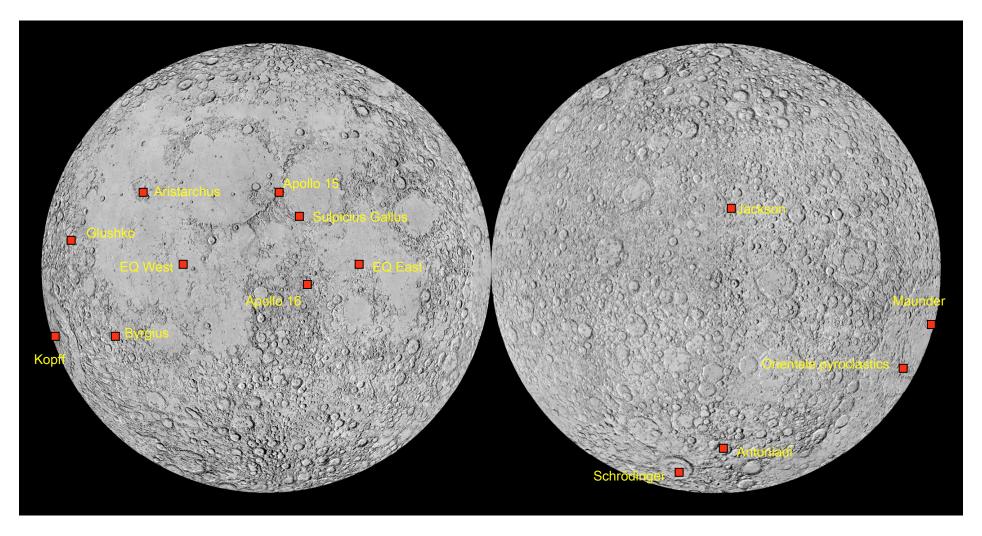








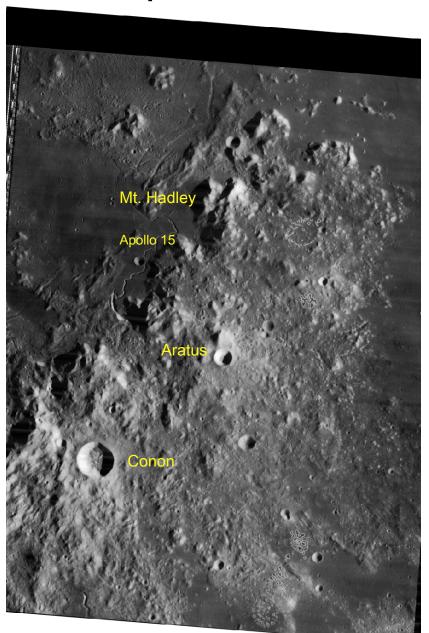








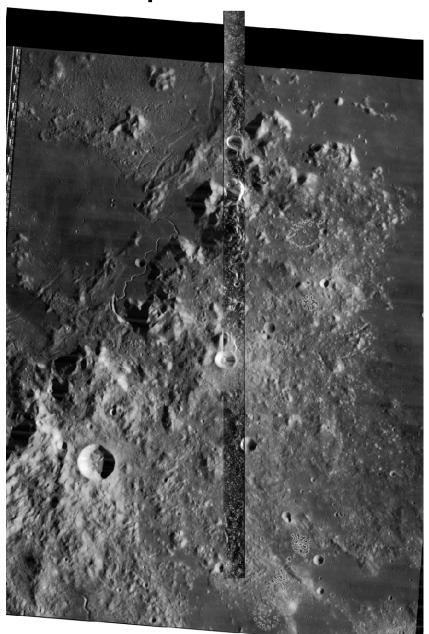








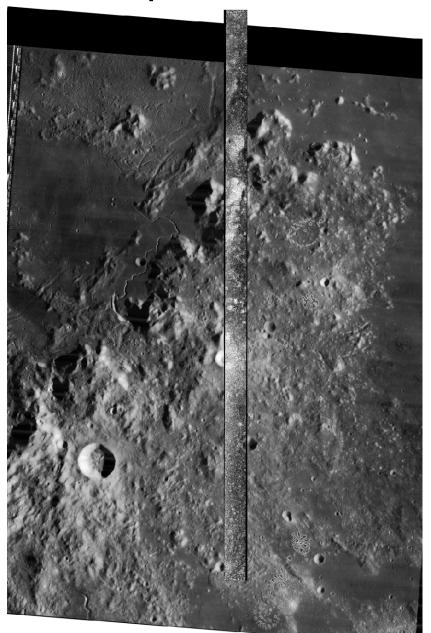








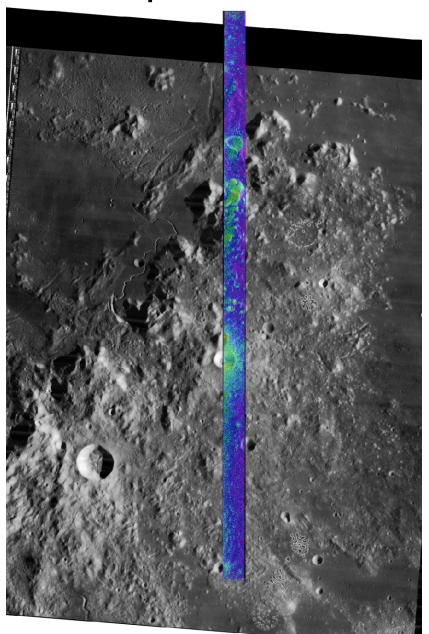








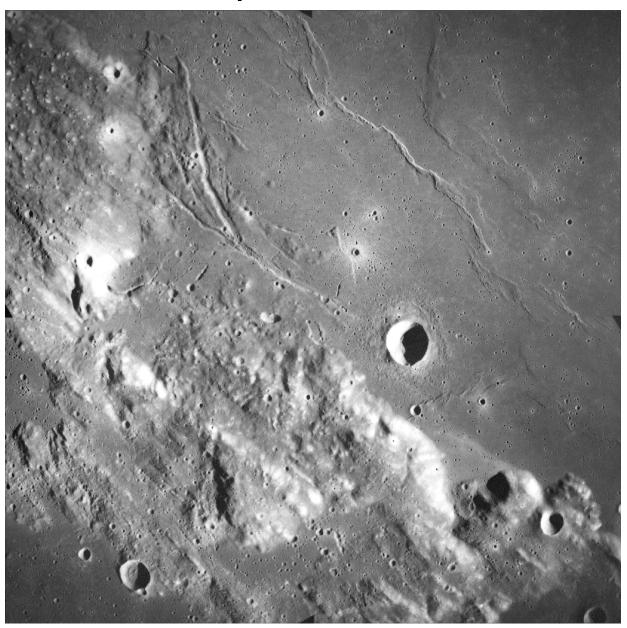








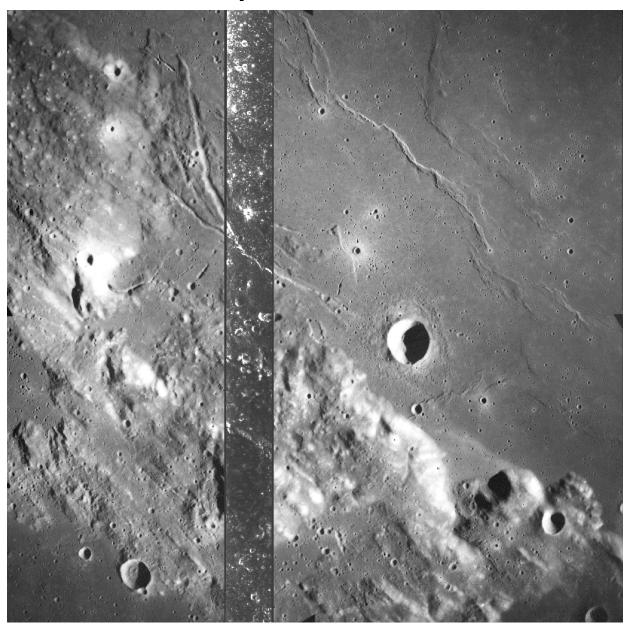








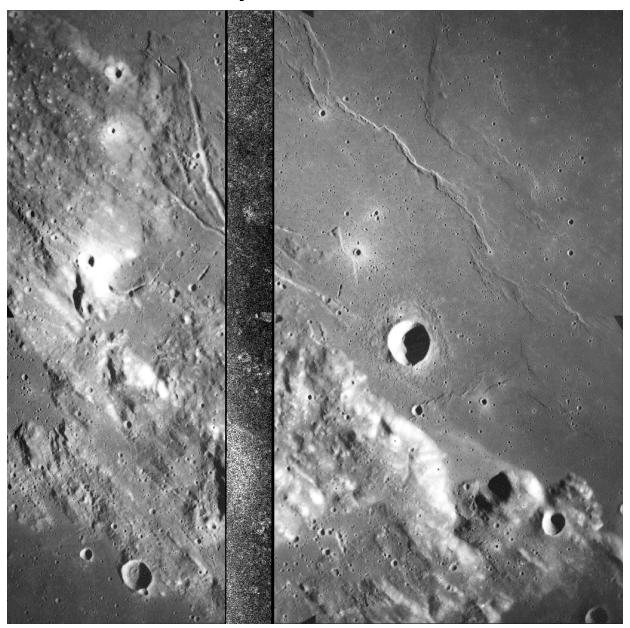








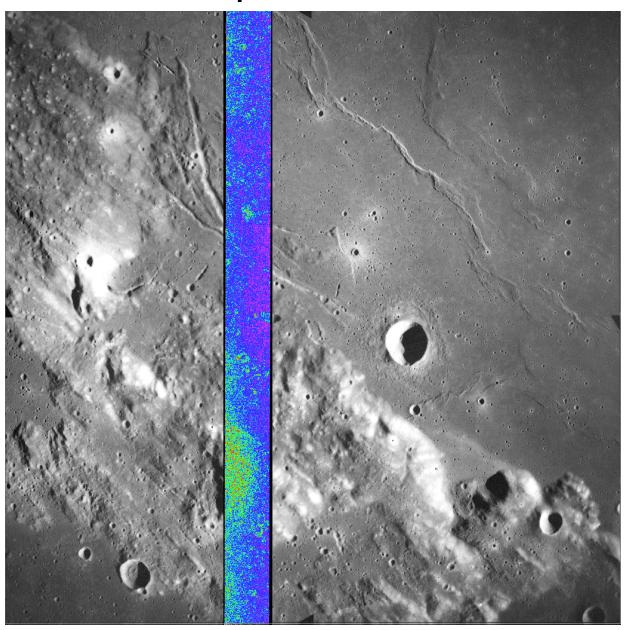










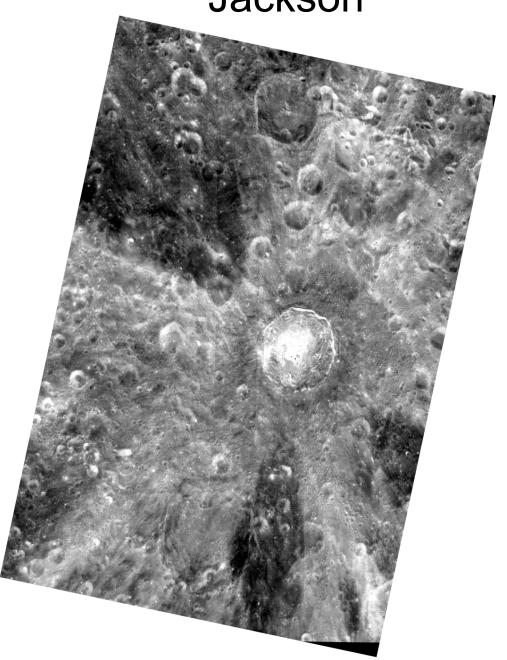






## Jackson



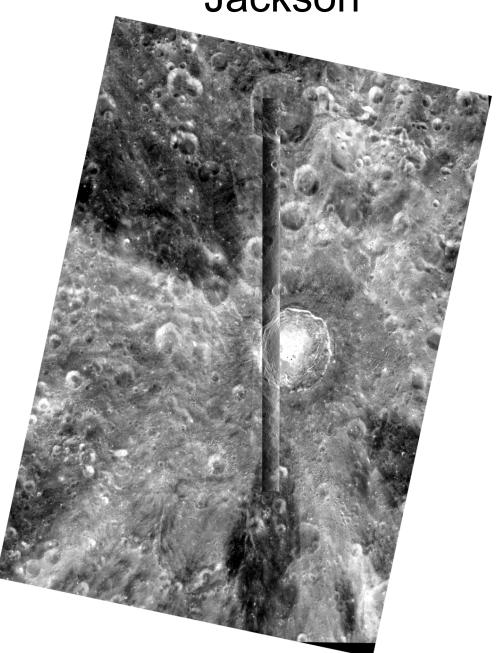






## Jackson



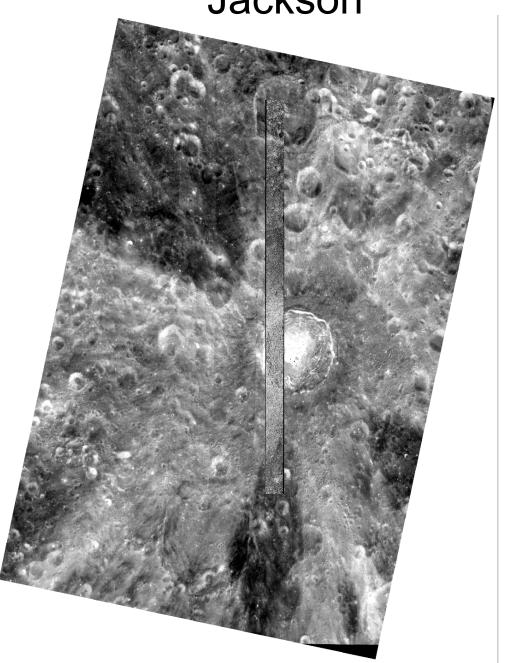










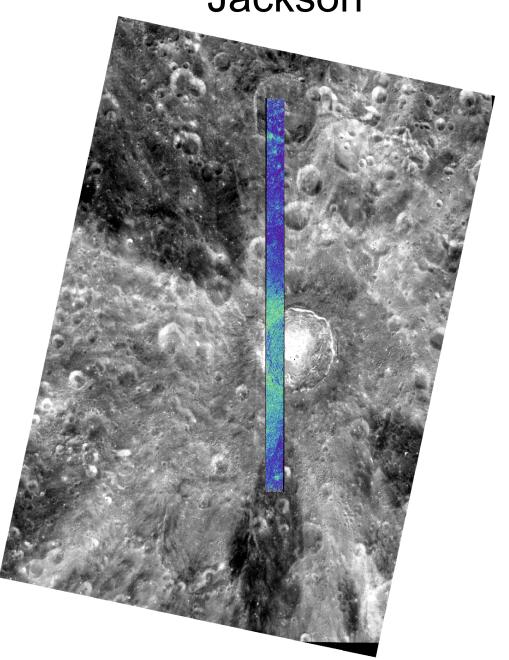






## Jackson





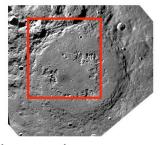








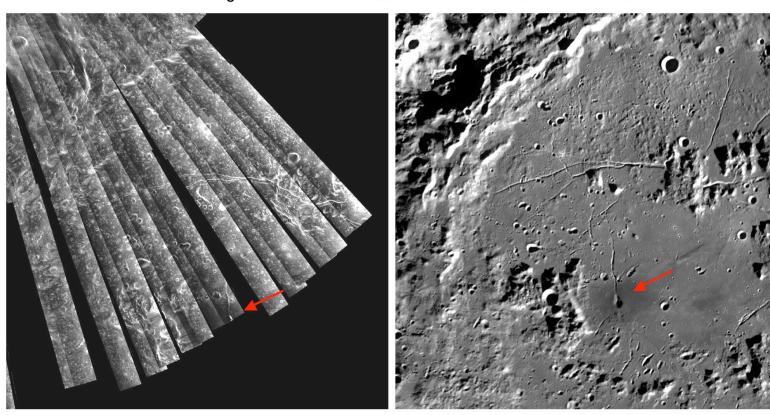
#### Schrödinger Basin Lunar far side 320 km diameter; Center 75° S, 132° E



Mini-SAR images

10 km

Clementine mosaics

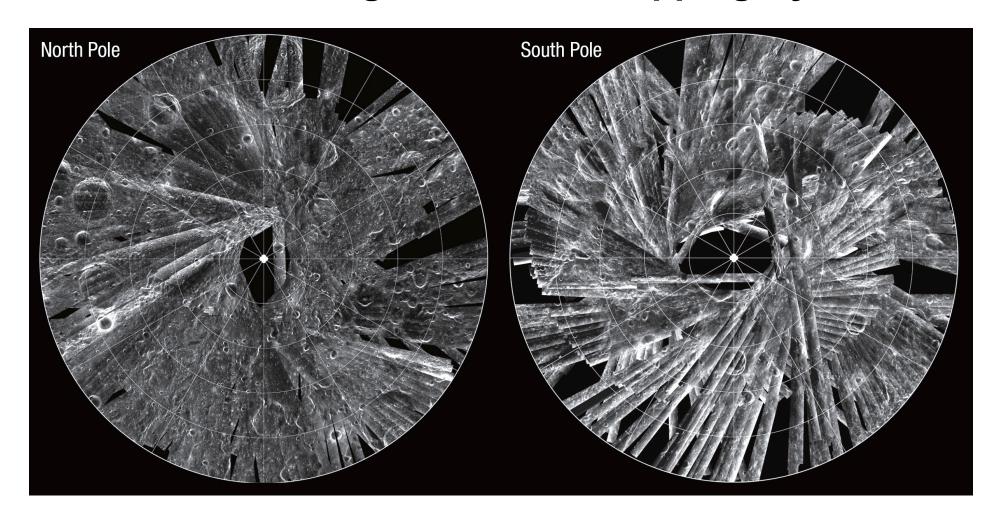








## **MSR Coverage from First Mapping Cycle**

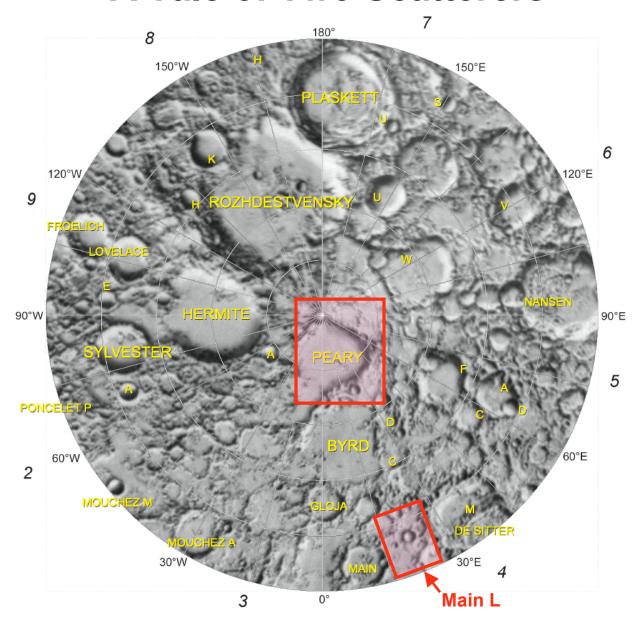






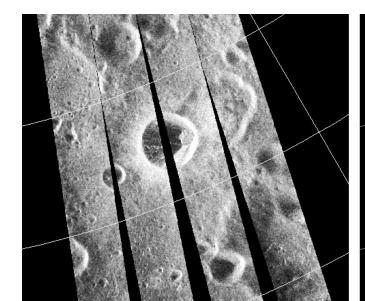


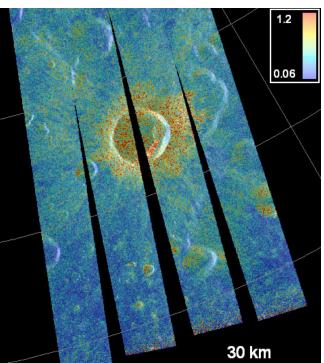
### **A Tale of Two Scatterers**









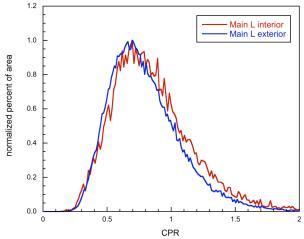


OS SAR image

Main L 14 km diameter 81.4° N, 22° E

30 km

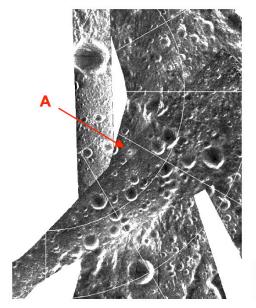
CPR image



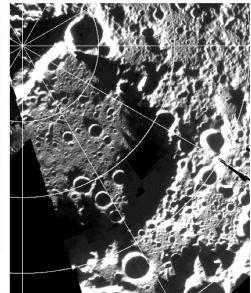








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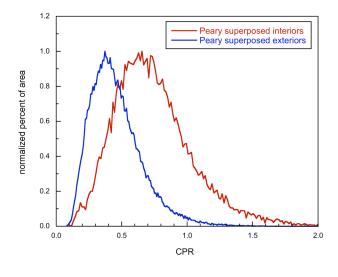


OS SAR mosaic

CPR mosaic

Floor of Peary 73 km diameter 88.6° N, 33° E

Clementine hires mosaic

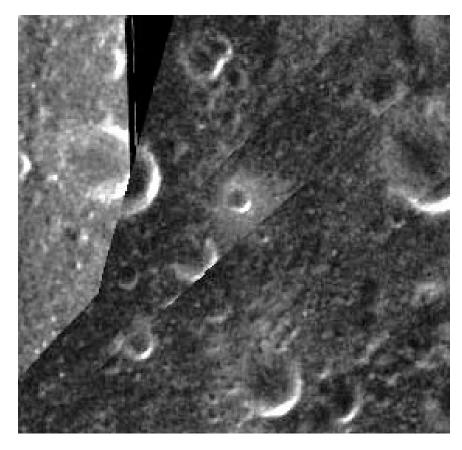








### **Close-up of Floor of Peary**



OS SAR image

CPR SAR image





### Summary



Mini-SAR is operational and is mapping the poles of the Moon Calibration data has been obtained and SAR data is being reprocessed using this information

We have mapped about 95% of both polar areas; due to some operational issues, coverage is not contiguous

Areas of high CPR have been identified:

Some high CPR is clearly associated with surface roughness (e.g., Main L ejecta blanket)

Some deposits (e.g., near north pole on floor of Peary) show high CPR and are restricted to the *interior* of craters; these features are in permanent darkness. Cause under investigation

#### **Future Work:**

Next mapping cycle mid-August to mid-October, 2009 (from 200 km orbit)

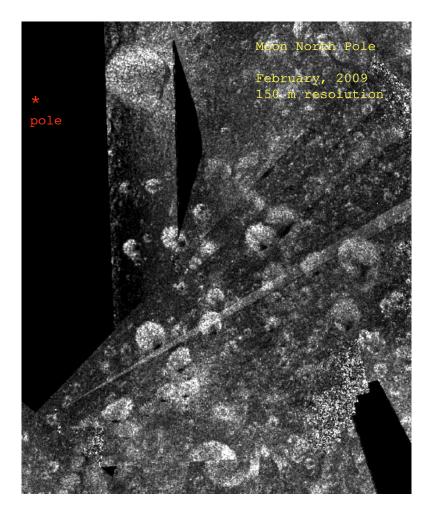
Complete mapping of both poles in both look directions
Fill-in polar (and other) gaps in coverage (SAR and scatterometry)
Conduct Stokes analysis of polar deposit backscatter properties
Compare and correlate with other data (e.g., topo, neutron, thermal)

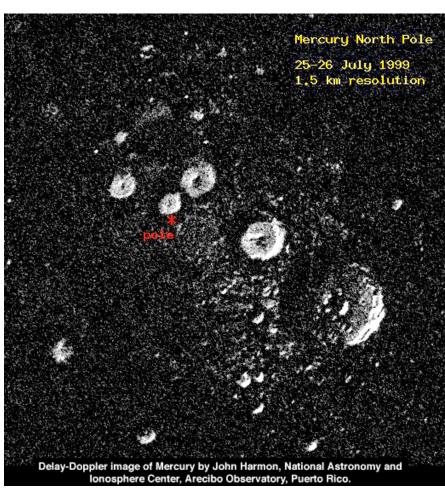






### **High CPR at North Pole**





Moon Mercury

